

**XXVI International Physics Olympiad**

**Canberra, ACT  
Australia**

**Theoretical Competition**

**July 7, 1995**

**Time Allowed: 5 Hours**

**READ THIS FIRST**

**Permitted Materials: Non Programable Calculators**

**Instructions:**

- 1. Use only the pen provided**
- 2. Use only the marked side of the paper**
- 3. Begin each problem on a separate sheet**
- 4. Write at the top of every sheet:**
  - The number of the problem**
  - The number of the sheet in your solution for each problem**
  - The total number of sheets in your solution to the problem.**

**Example (for Problem 1 with 3 sheets):** 1    1/3

1    2/3  
1    3/3

**Do not staple your sheets. They will be clipped together for you at the end of the examination.**

## Question 1

### Gravitational Red Shift and the Measurement of Stellar Mass

(a) (3 marks)

A photon of frequency  $f$  possesses an effective inertial mass  $m$  determined by its energy. We may assume that it has a gravitational mass equal to this inertial mass. Accordingly, a photon emitted at the surface of a star will lose energy when it escapes from the star's gravitational field. Show that the frequency shift  $\Delta f$  of a photon when it escapes from the surface of the star to infinity is given by

$$\frac{\Delta f}{f} \approx -\frac{GM}{Rc^2}$$

for  $\Delta f \ll f$  where

$G$	=	gravitational constant
$R$	=	radius of the star
$c$	=	velocity of light
$M$	=	mass of the star.

Thus, the red shift of a known spectral line measured a long way from the star can be used to measure the ratio  $M/R$ . Knowledge of  $R$  will allow the mass of the star to be determined.

(b) (12 marks)

An unmanned spacecraft is launched in an experiment to measure both the mass  $M$  and radius  $R$  of a star in our galaxy. As the spacecraft approaches its objective radially, photons emitted from  $\text{He}^+$  ions on the surface of the star are monitored via resonance excitation of a beam of  $\text{He}^+$  ions in a test chamber inside the spacecraft. Resonance absorption occurs only if the  $\text{He}^+$  ions are given a velocity towards the star to allow exactly for the red shifts. The velocity ( $v = \beta c$ ) of the  $\text{He}^+$  ions in the spacecraft relative to the star at absorption resonance is measured as a function of the distance  $d$  from the (nearest) surface of the star. The experimental data are displayed in the accompanying table. Fully utilize the data to determine graphically the mass  $M$  and radius  $R$  of the star. There is no need to estimate the uncertainties in your answer.

#### Data for Resonance Condition

Velocity parameter	$\beta = \frac{v}{c}$ ( $10^{-5}$ )	3.352	3.279	3.195	3.077	2.955
Distance from surface of star	$d$ ( $10^8\text{m}$ )	38.90	19.98	13.32	8.99	6.67

(c) In order to determine  $R$  and  $M$  in such an experiment, it is usual to consider the frequency correction due to the recoil of the emitting atom. [Thermal motion causes emission lines to be broadened without displacing distribution maxima, and we may therefore assume that all thermal effects have been taken into account.]

(i) (4 marks)

Let  $E$  be the energy difference between two atomic energy levels, with the atom at rest in each case. Assume that the atom decays at rest, producing a photon and a recoiling atom. Obtain the relativistic expression for the energy  $hf$  of a photon emitted in terms of  $E$  and the initial rest mass  $m_0$  of the atom.

(ii) (1 mark)

Hence, make a numerical estimate of the relativistic frequency shift  $\left(\frac{\Delta f}{f}\right)_{\text{recoil}}$

for the case of  $\text{He}^+$  ions.

Your answer should turn out to be much smaller than the gravitational red shift obtained in part (b).

Data:

$$\text{Velocity of light } c = 3.0 \times 10^8 \text{ m/s}$$

$$\text{Rest energy of He } m_0 c^2 = 4 \times 938 \text{ (MeV)}$$

$$\text{Bohr energy } E_n = -\frac{13.6 Z^2}{n^2} \text{ (eV)}$$

$$\text{Gravitational constant } G = 6.7 \times 10^{-11} \text{ Nm}^2 \text{ kg}^{-2}.$$